

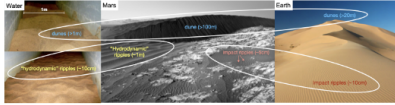
# COEVOLVING AERODYNAMIC AND IMPACT RIPPLES ON EARTH: UNIFYING BEDFORMS ON WATER, EARTH AND MARS

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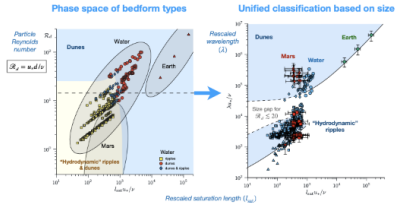
## 1. What we think we know

This is current understanding of bedform types in unidirectional flows [2, 1]:

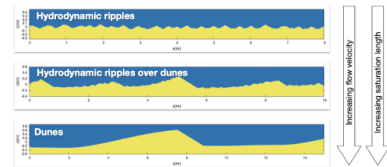


where hydrodynamic ripples are defined as the smallest bedforms of hydro(aero)dynamic origin, in contrast to the so-called impact ripples which arise from the interaction of hopping grains with the bed surface.

**Theory** – This classification was based on an improved understanding of the underlying physics. In agreement with the morphodynamic model [1], existing bedform data can be collapsed after rescaling all relevant lengths (grain size  $d$ , the transport saturation length  $L_{sat}$  and bedform wavelength  $\lambda$ ) by the viscous length  $\nu/\omega_*$ , where  $\nu$  is the kinematic viscosity of the fluid and  $\omega_*$  is the friction speed in the fluid boundary layer. The transport saturation length  $L_{sat}$  is defined as the spatial lag between the bed shear stress and the transport rate.

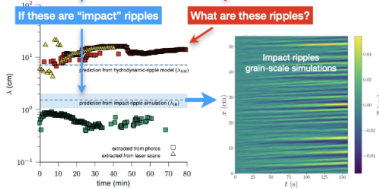
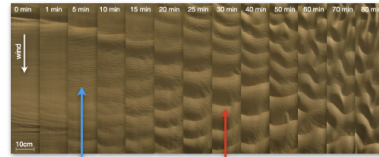


Here are examples of the bedform types for water using morphodynamic simulations:



## 2. New unexplained data

Recent wind tunnel experiments showed the co-evolution of two different small bedforms for fine spherical grains ( $d \approx 90\mu\text{m}$ ) on Earth [3], instead of the expected single impact ripple.



## 5. Open questions

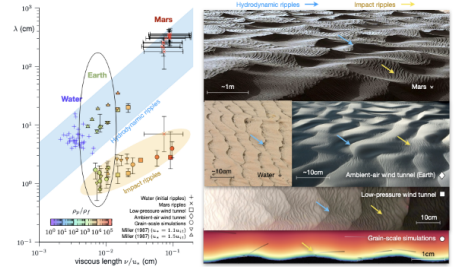
- Why most ripples on the field (unimodal sand) only seem to have a single length scale?
- Are ripples in the field "impact" or "hydrodynamic" ones?
- What is the role of grain size and shape, and flow velocity?
- What is the physical origin of the smaller saturation length?
- Is there a smaller saturation length on Mars as well?

## References

[1] Orencio Duran Vincent et al. "A Unified Model of Ripples and Dunes in Water and Planetary Environments". In: Nature Geoscience 12.5 (2019).  
 [2] M. G. A. Lapotre et al. "Large Wind Ripples on Mars: A Record of Atmospheric Evolution". In: Science 353.6294 (2016), pp. 55–58.  
 [3] Hezi Yizhaq et al. "Coevolving Aerodynamic and Impact Ripples on Earth". In: Nature Geoscience (2024), pp. 1–7. DOI: 10.1038/s41561-023-01348-3.

## 3. Interpretation as hydrodynamic ripples on Earth

A comparison with existing data for water and Mars, as well as new impact ripples simulations and new low-pressure wind tunnel experiments (for fine sand and flow velocities close to the transport threshold), suggests the larger ripple can be interpreted as a hydrodynamic ripple, clearly distinct from the new smaller impact ripple [3].



## 4. Proposed physical mechanism

But hydrodynamic ripples ( $\lambda \sim 0.1\text{m}$ ) should not exist on Earth because  $L_{sat} \sim 0.5\text{m}$ ; and by definition  $\lambda > L_{sat}$ . However, new evidence suggests the existence of a much smaller saturation length  $\sim 1\text{cm}$  (left panel), that can explain the data using linear instability analysis (right panel) [3].

